## STELLAR ABUNDANCES in CLUSTERS

High-resolution, multi-object spectroscopy of globular \& open clusters in the Milky Way

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## There are more surveys in heaven and earth, Horatio...

- large scale

Gaia-ESO public spectroscopic survey
P. Donati, D. Romano, M. Tosi, L. Magrini, S. Randich,
T. Cantat Gaudin, R. Sordo, A. Vallenari, E. Friel, H. Jacobson \& GES Consortium (400+)

- "private"

FLAMES GC program
E.Carretta, V. D'Orazi, R. Gratton, S. Lucatello,
A. Sollima, C. Sneden

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Way and Local Group Galaxies" (PI F. Matteucci)


- "private"


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## The Gaia-ESO Survey in a nutshell



For information : http://www.gaia-eso.eu

- PI Randich/Gilmore
- 400+ researchers
- 300 VLT nights / 5 years (3 yrs / 32 obs runs)
- FLAMES
- $10^{5} \mathrm{MW}$ stars
- 70+ open clusters
(~30 observed)
- STD / GCs
- distributed analysis


# The Gaia-ESO Survey - open clusters 

## core science

- Open cluster formation \& dynamics - Properties \& evolution of MW disk
- Stellar evolution


## The Gaia-ESO Survey - all OCs

- 18 months
- 24 months
- 31 months
$\square$ protected
courtesy
S. Randich



## The Gaia-ESO Survey - old OCs




## The Gaia-ESO Survey - old OCs



Yong+2012 \& literature [Fe/H] after 2012 ; BOCCE age,Rgc if available

## The Gaia-ESO Survey - Trumpler 20



Carraro+2010


GES - Donati+ 2014

## The Gaia-ESO Survey - Trumpler 20




GES - Donati+ 2014 : ~40\% M

## The Gaia-ESO Survey - Trumpler 20




## The Gaia-ESO Survey - inner disk OCs

| OC | age | Rgc | RV | [ $\mathrm{Fe} / \mathrm{H}$ ] | GES paper |
| :---: | :---: | :---: | :---: | :---: | :---: |
| Tr 20 | 1.5 | 6.9 | -40.5 (1470) | +0.16 (13) | Donati+2014 |
| NGC4815 | 0.5 | 6.9 | -30.2 (218) | +0.03 (5) | Friel+2014 |
| NGC6705 | 0.3 | 6.3 | +34.5 (1053) | +0.10 (21) | Cantat-Gaudin+2014 |
| Be 81 | 1.0 | 5.5 | +47.6 (280) | in progress | in progress |
| NGC6208 | 0.7 | 7.4 |  |  | observed |
| Be 44 | 1.3 | 7.1 |  |  | observed |
| Pismis 18 | 1.2 | 6.8 |  |  | observed |
| Tr 23 | 1.0 | 6.2 |  |  | observed |
| NGC6005 | 1.2 | 6.0 |  |  | observed |
| RV: GIRAFFE+UVES - [FE/H]: UVES |  |  |  |  |  |

## for more details, see talk by Laura Magrini (on Tuesday)

## The Gaia-ESO Survey - 3 inner disk OCs



__ chemev models Romano+2010
Trumpler 20
NGC4815
GES - Magrini+ 2014

## The Gaia-ESO Survey - 3 inner disk OCs





## End of first act ...

## Open clusters :

> information on Galactic disk

> e.g. structure
> e.g. chemical evolution
> test of stellar models

$$
\begin{aligned}
& \text { e.g., tracks } \\
& \text { e.g., mixing mechanisms }
\end{aligned}
$$

> homogenous clusters (single populations)

## GCs: setting the stage

In GCs light elements show
star-to-star variations
giants

ARA\&A 2004
$\sim 400$ stars
$\sim 20 \mathrm{GCs}$

In field stars they do not


Field stars (Gratton+ 2000)

## Our sample of clusters

- 24 massive GCs : Mv=-5.5 to -10





Piotto+2002,
HST snapshot

NGC2808

NGC6388

## GCs: $\mathrm{Na} \& \mathrm{O}$ in RGB stars



Carretta+ 2009,2010, 2011,2013, 2014, ...

## GCs: $\mathrm{Na} \& \mathrm{O}$ in RGB stars



## GCs: $\mathrm{Na} \& \mathrm{O}$ in RGB stars



FG: P
SG:I+E



## GCs: $\mathrm{Na} \& \mathrm{O}$ in RGB stars




Carretta+2009, 2010

## GCs: $\mathrm{Na} \& \mathrm{O}$ in RGB stars




## All creatures great and small. I. GCs



some $10^{6}$
mass in $\mathbf{M}_{\odot}$

some $10^{4}$

GCs with $\mathrm{Na}-\mathrm{O}$ anticorrelation
Carretta 2010 etc

## All creatures great and small. I. GCs


$\diamond$ Ter 7, Pal 12 : no?
$\diamond$ Rup 106 : noTerzan 8 : yes?Be39, NGC 6791 : no

- NGC 6535, NGC 6139 : ??

Ter7:Tautvaisiene+,Sbordone+
Pal12: Coben
Rup106: Villanova+

## All creatures great and small. I. GCs



FG dominates

Carretta+2014


## All creatures great and small. II. OCs



MacLean, De Silva, Lattanzio 2015

## All creatures great and small. II. OCs





## All creatures great and small. II. OCs



Hydra@WIYN
HIRES@Keck

NGC 6791:
$\mathrm{Mv}=-4.14$; age=8 Gyr
$[\mathrm{Fe} / \mathrm{H}]=+0.4$

## All creatures great and small. II. OCs

NGC 6791


## All creatures great and small. II. OCs



## All creatures great and small. II. OCs



## All creatures great and small. II. OCs





## Results \& legacy value

$>$ clusters are halo and disk (and bulge) tracers
$>$ constraints for stellar models
$>$ cover both young \& old, metal-poor \& rich populations
$>$ combine photometry, spectroscopy \& models
... wait for Gaia-ESO, APOGEE, etc and - of course - Gaia ...

## IAU Symposium 317 at the 2015 IAU GA


http://www.iau.org/science/meetings/future/symposia/1124/
http://www.eso.org/~marnabol/IAU317_index.html

