STELLAR ABUNDANCES in CLUSTERS

High-resolution, multi-object spectroscopy of globular & open clusters in the Milky Way

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There are more surveys in heaven and earth, Horatio...

• large scale

Gaia-ESO public spectroscopic survey

P. Donati, D. Romano, M. Tosi, L. Magrini, S. Randich,

T. Cantat Gaudin, R. Sordo, A. Vallenari, E. Friel, H. Jacobson
 & GES Consortium (400+)

• "private"

FLAMES GC program

E.Carretta, V. D'Orazi, R. Gratton, S. Lucatello,

A. Sollima, C. Sneden

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PRIN-MIUR 2010-11 :

"The chemical and Dynamical Evolution of the Milky

Way and Local Group Galaxies" (PI F. Matteucci)

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The Gaia-ESO Survey in a nutshell



- PI Randich/Gilmore
- 400+ researchers
- 300 VLT nights/5 years (3 yrs /32 obs runs)
- FLAMES
- 10⁵ MW stars
- 70+ open clusters (~30 observed)
- STD / GCs
- distributed analysis



For information : http://www.gaia-eso.eu

The Gaia-ESO Survey – open clusters

core science

- Open cluster formation & dynamics
- Properties & evolution of MW disk
- Stellar evolution

The Gaia-ESO Survey – all OCs



The Gaia-ESO Survey – old OCs



The Gaia-ESO Survey – old OCs



Yong+2012 & literature [Fe/H] after 2012; BOCCE age, Rgc if available

The Gaia-ESO Survey – Trumpler 20





The Gaia-ESO Survey – Trumpler 20



GES - Donati+ 2014 : ~40% M

Chemical and dynamical evolution of the Milky Way and Local Group galaxies

The Gaia-ESO Survey – Trumpler 20





GES - Donati+ 2014

The Gaia-ESO Survey – inner disk OCs

| OC | age | Rgc | RV | [Fe/H] | GES paper |
|--------------|-----|-----|--------------|-------------|--------------------|
| Tr 20 | 1.5 | 6.9 | -40.5 (1470) | +0.16 (13) | Donati+2014 |
| NGC4815 | 0.5 | 6.9 | -30.2 (218) | +0.03 (5) | Friel+2014 |
| NGC6705 | 0.3 | 6.3 | +34.5 (1053) | +0.10 (21) | Cantat-Gaudin+2014 |
| Be 81 | 1.0 | 5.5 | +47.6 (280) | in progress | in progress |
| NGC6208 | 0.7 | 7.4 | | | observed |
| Be 44 | 1.3 | 7.1 | | | observed |
| Pismis 18 | 1.2 | 6.8 | | | observed |
| <i>Tr 23</i> | 1.0 | 6.2 | | | observed |
| NGC6005 | 1.2 | 6.0 | | | observed |

RV: **GIRAFFE+UVES** - **[FE/H]**: **UVES**

for more details, see talk by Laura Magrini (on Tuesday)

The Gaia-ESO Survey – 3 inner disk OCs



The Gaia-ESO Survey – 3 inner disk OCs



End of first act ...

Open clusters :

information on Galactic disk

e.g. structure e.g. chemical evolution

test of stellar models

e.g., tracks

e.g., mixing mechanisms

homogenous clusters (single populations)



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Our sample of clusters 24 massive GCs : Mv=-5.5 to -10





Carretta+ 2009,2010, 2011,2013, 2014, ...





0





Carretta+2013

IQR=Inter Quartile Range= middle 50%





GCs with Na-O anticorrelation



some 10⁶ mass in M_⊙

some 10⁴

Carretta+2010 etc



GCs with Na-O anticorrelation

♦ Ter 7, Pal 12 : no?

- - Terzan 8 : yes?
- **Be39, NGC 6791**: no

Ter7 : Tautvaisiene+,Sbordone+ Pal12: Cohen Rup106 : Villanova+





MacLean, De Silva, Lattanzio 2015



Berkeley 39: Mv=-4.28 ; age=6 Gyr [Fe/H]=-0.2



Bragaglia+2012



NGC 6791: Mv=-4.14 ; age=8 Gyr [Fe/H]=+0.4

Bragaglia, Sneden, Gratton et al. 2014

NGC 6791

All creatures great and small. II. OCs

1.0 0.8 **Relative Flux** [0 I] 0.6 0.4 best CN-0.6 0.2 best CN best CN+0.6 0.0 6300 6302 Nal Cal Cal NaI Cal Cal 1.00 0.50 best Na,Ca - 0.5 best Na.Ca best Na,Ca + 0.5 star 15592 (a) (b) 0.00 6154 6156 6160 6162 6164







Results & legacy value

- clusters are halo and disk (and bulge) tracers
- constraints for stellar models
- > cover both young & old, metal-poor & rich populations
- combine photometry, spectroscopy & models

... wait for Gaia-ESO, APOGEE, etc and - of course - Gaia ...

IAU Symposium 317 at the 2015 IAU GA

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